

Overview of the Yambo code: main features and performance

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I. The yambo code

II. abintio MBPT & the QP concept

III. Excitons: the BSE



Part I The yambo code





www.yambo-code.eu

MaX flagship code





Properties

GPL

Quasi-particles Optics and excitons Magneto-optics & dichroism Electron-phonon coupling Real-time propagation Non-linear optics

Development (pre-GPL) version

Exciton-phonon coupling Pump and probe experiments Defects withing MBPT Ehrenfest dynamics Magnons

Applications









the ambo team



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Online documentation and tutorials

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team



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Community & Publications

Growing community of users using Yambo for forefront research. More than 200 publications.

Perfomances

GW study of Graphene @ Co(0001) interface



Yambo compiled with ifort (intel). **MPI + OpenMP**

mpirun -np #MPI #MPI=4 #THREADS=24 (2*#cores / #MPI)

Juwels-Cluster. 48 Intel cores per node



mpirun -np #MPI #MPI=4 (= #cards per node) #THREADS=8 (no effect here)

Juwels-Booster. 48 AMD cores per node 4 **Nvidia** A100 cards per node





Data available at: http://www.gitlab.com/max-centre/Benchmarks



Part II Abinitio Many-body Perturbation-Theory

Many-Body Perturbation-Theory

$$H = \sum_{i=1}^{N} h(\mathbf{r}_i) + \sum_{\substack{i,j=1\\i\neq j}}^{N} V(\mathbf{r}_i, \mathbf{r}_j)$$

1 - Define the Green function (GF)

$$G\left(\mathbf{r}',t';\mathbf{r},t\right) = -i\left\langle\phi_{0}^{N}\left|\hat{T}\left[\hat{\psi}\left(\mathbf{r}',t'\right)\hat{\psi}^{\dagger}(\mathbf{r},t)\right]\phi_{0}^{N}\right.\right\rangle$$



Interacting particles Ground state

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Interacting particles Ground state



Why the GF?

Leehman's representation

$$G(\mathbf{r}, \mathbf{r}', \omega) = \lim_{\eta \to 0^+} \sum_{n} \left\{ \frac{\chi_0^{N+1}(\mathbf{r}) \left[\chi_0^{N+1} \left(\mathbf{r}'\right)\right]^*}{\omega - \epsilon_n (N+1) - \mu + i\eta} + \frac{\left[\chi_0^{N-1} \left(\mathbf{r}\right)\right]^* \chi_0^{N-1}(\mathbf{r}')}{\omega - \epsilon_n (N-1) - \mu - i\eta} \right\}$$

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Phys. Rev. B 92, 125440 (2015)

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Excitation energies

$$\epsilon_n(N \pm 1) = E_n(N \pm 1) - E_0(N \pm 1)$$

Chemical potential

$$\mu = \begin{cases} E_0(N+1) - E_0(N) \\ E_0(N) - E_0(N-1) \end{cases}$$



Phys. Rev. B 92, 125440 (2015)

The QP concept





$G(\omega) = G^{0}(\omega) + G^{0}(\omega) \Sigma^{xc}[G](\omega)G(\omega)$

Interacting particles



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#poles of G >> #poles of G⁰

The QP concept



DFT vs MBPT



DFT + MBPT



G. Onida, L. Reining, and A. Rubio, Rev. Mod. Phys. 74, 601 (2002)

theVa

DFT $\left[\frac{-\nabla^2}{2} + v^{ext} + v^{Hxc}\right] \psi_{nk}(r) = \epsilon_{nk} \psi_{nk}(r)$

MBPT

$$G^{KS}(r,r',\omega) = \sum_{nk} \frac{\psi_{nk}^*(r)\psi_{nk}(r')}{\omega - \epsilon_{nk}^{KS} + i\eta}$$

$$G = G^{KS} + G^{KS} (\Sigma^{xc} - v^{xc}) G$$

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MBPT



abinitio **MBPT**







theYambo team



Weakly interacting quasi-particles



DFT and GW bands

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DFT vs GW band structure in hBN Photoemission spectral function 10 Energy relative to VBM (eV) 49999⁹⁹ 9999 -20 KT, K H M Г ω εsat $\epsilon^{\rm qp}$ ε μ Phys, Rev. Lett. 96, 026402 (2006)

theVam

See more in the next talk

QP complex energies



Adapted from M. Gatti PhD thesis

QP complex energies



The role of screening



$$W^{RPA} = v + v \chi^{RPA} v$$

$$\chi^{RPA}_{GG'}(q\,\omega)$$



$$\chi^{RPA} = \chi_0 + \chi_0 v \chi^{RPA}$$



Part III Excitons: The Bethe-Salpeter equation

Neutral excitations

1 - Define a two body response function

 $G(1,2,3,4) = (-i)^2 \langle N | T [\psi(1)\psi(3)\psi^{\dagger}(4)\psi^{\dagger}(2)] | N \rangle \qquad 1 = (x_1, t_1, \sigma_1)$

 $L(1,2,3,4) = L^{0}(1,2,3,4) - G(1,2,3,4)$

Strongly interacting (quasi)electron – (quasi)hole

Neutral excitations

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 $G(1,2,3,4) = (-i)^2 \langle N | T [\psi(1)\psi(3)\psi^{\dagger}(4)\psi^{\dagger}(2)] | N \rangle \qquad 1 = (x_1, t_1, \sigma_1)$

$$L(1, 2, 3, 4) = L^{0}(1, 2, 3, 4) - G(1, 2, 3, 4)$$

2- Define a Kernel

$$K^{Hxc}[G](1,2;3,4) = \frac{\delta \Sigma(1,2)}{\delta G(3,4)}$$

Strongly interacting (quasi)electron – (quasi)hole



Neutral excitations

1 - Define a two body response function

 $G(1,2,3,4) = (-i)^2 \langle N | T [\psi(1)\psi(3)\psi^{\dagger}(4)\psi^{\dagger}(2)] | N \rangle \qquad 1 = (x_1, t_1, \sigma_1)$

$$L(1, 2, 3, 4) = L^{0}(1, 2, 3, 4) - G(1, 2, 3, 4)$$

Strongly interacting (quasi)electron – (quasi)hole



$$K^{Hxc}[G](1,2;3,4) = \frac{\delta \Sigma(1,2)}{\delta G(3,4)}$$



3- Spectral representation $(t_1 = t_2)$ $(t_3 = t_4)$ $L(\omega) \longrightarrow \text{poles } \omega_I = E_I(N) - E_0(N)$

the

Bethe-Salpeter Equation

The Dyson equation for L, or Bethe Salpeter Equation (BSE)

 $L = L^{qp} + L^{qp} K^{Hxc} L$

$$K^{Hxc}(\omega=0)=(v-W)$$

can be rewritten as an eigenvalue problem

$$\left[\left(\epsilon_{ck}-\epsilon_{vk-q}\right)+v_{cvk,c'v'k'}-W_{cvk,c'v'k'}\right]A_{c'v'k}^{\lambda q}=\omega_{\lambda q}A_{cvk}^{\lambda q}$$

Bethe-Salpeter Equation

The Dyson equation for L, or Bethe Salpeter Equation (BSE)

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Excitation wave-function

can be rewritten as an eigenvalue problem

$$\left[\left(\epsilon_{ck}-\epsilon_{vk-q}\right)+v_{cvk,c'v'k'}-W_{cvk,c'v'k'}\right]A_{c'v'k'}^{\lambda q}=\omega_{\lambda q}A_{cvk}^{\lambda q}$$

$$\begin{aligned} \left|\Psi^{\lambda q}(k)\right|^{2} &= \sum_{cv} \left|A_{cvk}^{\lambda q}\right|^{2} \\ \Psi^{\lambda q}(x_{h}, x_{e}) &= \sum_{cvk} A_{cvk}^{\lambda q} \psi_{ck-q}^{*}(x_{e}) \psi_{vk}(x_{h}) \end{aligned}$$

abinitio GW+BSE



BSE overview



Exciton wave-function





ACTIVE DEVELOPERS (alphabetic order)

- * Ignacio Martin Alliati
- * Claudio Attaccalite
- * Miki Bonacci
- * Elena Cannuccia
- * Andrea Ferretti
- * Myrta Gruening
- * Alberto Guandalini
- * Conor Hogan
- * Dario Alejandro Leon-Valido
- * Andrea Marini
- * Alejandro Molina-Sánchez
- * Fulvio Paleari
- * Maurizia Palummo
- * Davide Sangalli
- * Nicola Spallanzani
- * Daniele Varsano

FORMER DEVELOPERS

- * Fabio Affinito
- * David Kammerlader
- * Ivan Marri
- * Antimo Marrazzo
- * Margherita Marsili
- * Pedro Melo
- * Henrique Miranda
- * Ryan McMillan



Thank you for your attention

- Many-body perturbation theory calculations using the yambo code Journal of Physics: Condensed Matter 31, 325902 (2019)
- 2. Yambo: an ab initio tool for excited state calculations Comp. Phys. Comm. 144, 180 (2009)

More slides

- 1. Many-body perturbation theory calculations using the yambo code Journal of Physics: Condensed Matter 31, 325902 (2019)
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How the code works



Yambo parallel performance



heterogeneous architectures: MPI + OpenMP + CUDA





optimisation of MPI+OpenMP parallelism

 working at scale (bottleneck identification and solution)



system size: 72+1 atoms, 2000 bands, 6 Ry for Xo repr (N=1317); ~290 occ states, 8 kpts.

data available at: <u>http://www.gitlab.com/</u> max-centre/Benchmarks

Yambo: performance (GPU)



heterogeneous architectures: MPI + OpenMP + CUDA



x14 wrt SKL (10 nodes)



x1.4 wrt V100 (2 nodes)



complete GW workflow for defected TiO2 (rutile)

small system, stress test

I MPI task/GPU

data obtained on Marconi100, 4 V100 GPUs/node
and DGX arch, 8 A100 GPUs/node



system size: 72+1 atoms, 2000 bands, 6 Ry for Xo repr (N=1317); ~290 occ states, 8 kpts.

data available at: <u>http://www.gitlab.com/</u> max-centre/Benchmarks



Part IV More applications